Optical Engineering

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Frédéric Allegrini Robert W. Ebert Stephen A. Fuselier Georgios Nicolaou Peter Bedworth Steve Sinton Karlheinz J. Trattner



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Frédéric Allegrini, a,b,* Robert W. Ebert, Ebert, Ebert, Georgios Nicolaou, Ebert, Carben A. Fuselier, Georgios Nicolaou, Peter Bedworth, Steve Sinton, and Karlheinz J. Trattner J. Trattner J. Trattner J. Trattner J. Trattner Steve Sinton, Steve Sinton, Carben A. Fuselier, Abert Steve Sinton, Carben A. Fuselier, Carben A. Fuselier, Abert Steve Sinton, Carben A. Fuselier, Carben A. Fu

Abstract. Carbon foils have been used reliably for many decades in space plasma instrumentation to detect ions and energetic neutral atoms (ENA). When these particles pass through a foil, secondary electrons are emitted on both sides. Those electrons are used for coincidence detection and/or timing signals. Ultrathin carbon foils are also used to convert an ENA into an ion for further analysis. The interaction of particles with carbon foils also includes unwanted effects such as energy straggling and angular scattering, both of which scale with foil thickness. Therefore, it has always been a goal to use foils as thin as practically possible. The foils used in space are usually made of amorphous carbon of roughly a hundred atomic layers. Graphene can be made much thinner, even down to a single atomic layer, and is therefore a natural candidate for this kind of application. We evaluate one aspect of the interaction of particles with foils: charge exchange. We show the first measurements of exit charge state distributions of ~1 to 50 keV ions passing through self-supported graphene foils. We compare the charge state fraction of exiting particles with state-of-the-art amorphous carbon foils. © *The Authors. Published by SPIE under a Creative Commons Attribution 3.0 Unported License. Distribution or reproduction of this work in whole or in part requires full attribution of the original publication, including its DOI.* [DOI: 10.1117/1.0E.53.2.024101]

Keywords: graphene; carbon foil; space plasma instrument; energetic neutral atom; charge exchange.

Paper 131687 received Nov. 4, 2013; revised manuscript received Jan. 8, 2014; accepted for publication Jan. 13, 2014; published online Feb. 4, 2014.

1 Introduction

Carbon foils are used in space plasma instrumentation to enable the detection of ions and neutral atoms. These particles create secondary electron emission when passing through the foils.^{2,3} The secondary electrons are detected by electron multipliers-typically channel electron multipliers4 or microchannel plates (MCPs)⁵—to create a coincidence and/or a timing signal. Another result of this interaction is charge exchange, in which the projectile gains and loses electrons predominantly by interactions with the band structure of the foil. The exit charge state distribution mainly depends on the velocity and the atomic number of the particle. At keV energies, the incident charge state—the charge that a particle has before interacting with the foil—does not affect the exit charge state—the charge that a particle has after leaving the foil. Under these circumstances, the particle velocity is lower or comparable to the Bohr velocity (~26 keV/amu) so that foil conduction electrons are sufficiently mobile over the interaction time scales to freely interact with the projectile. Therefore, charge equilibrium of the particle is reached in the first few atomic layers of the foil and the charge state memory of the incident particle is lost.

The property of charge conversion with foils is used, for example, in different space plasma instruments—for solar wind, magnetospheric, and heliospheric missions—and as charge strippers in high-energy particle accelerators. Here, we focus only on space plasma instrumentation for ions (and neutral atoms) in the range from ~1 to 50 keV.

*Address all correspondence to: Frédéric Allegrini, E-mail: fallegrini@swri.edu

It is important to know the charge distribution upon exit from the foil. For example, the mass time-of-flight sensor part of the charge, element, and isotope analysis system (CELIAS) experiment⁶ on solar and heliospheric observatory (SoHO) has an isochronous time-of-flight section that uses a carbon foil. The ions of same mass-per-charge have an identical time-of-flight between the carbon foil and the detector. Therefore, it is imperative to know the exit charge state distribution of ions after the foil in order to calculate abundance ratios between species.

Another example where the exit charge state distribution is critical is the cassini plasma spectrometer/ion mass spectrometer (CAPS/IMS) sensor⁷ on Cassini. The particles, depending on their charge state, land on different detectors in the time-of-flight section. The charge state distributions are used to determine accurate mass spectra.

An extensive contribution to the literature on charge exchange at keV energies was done by the Space Research and Planetary Sciences Division at the University of Bern. They measured the charge state distributions after carbon foils of the major solar wind ions and proposed a mechanism to explain the charge exchange properties, which led to a semiempirical model. Other groups—e.g., the Space Plasma Physics Group at the Los Alamos National Laboratory $^{15-20}$ —also contributed significantly to the topic, especially for ultrathin ($\sim 0.5 \ \mu g/cm^2$ or $\sim 100 \ atoms$ thick) carbon foils.

The property of charge exchange is also critical to the detection of energetic neutral atoms (ENA). ENA imaging has become an important element of heliospheric research

^aSouthwest Research Institute, P.O. Drawer 28510, San Antonio, Texas 78228

^bUniversity of Texas at San Antonio, Physics and Astronomy Department, 1 UTSA Circle, San Antonio, Texas 78249

^cLockheed Martin Advanced Technology Center, 3251 Hanover Street, Palo Alto, California 94304

^dUniversity of Colorado, Laboratory for Atmospheric and Space Physics, 1234 Innovation Drive, Boulder, Colorado 80303

with the recent interstellar boundary explorer (IBEX)²¹ which has solidified this imaging technique as a critical component of future NASA heliospheric and magnetospheric missions.

Neutral atom imagers, such as those on IBEX, cover energies from 0.01 to 100 keV, encompassing the entire energy range of neutral atoms (and their parent ion populations) in the magnetosphere, heliosphere, and interstellar medium. To measure neutrals over this broad energy range, different types of imagers are required. Most imagers have a key subsystem that converts an incoming neutral atom into an ion. Then, the ion can be deflected away from the initial neutral direction. This deflection is critical for separating the original signal from much higher backgrounds (e.g., ultraviolet radiation) that are present in space. The conversion subsystem is the least efficient subsystem in a neutral atom imager, especially at energies below 1 keV. The state-of-the-art neutral atom camera IBEX-Hi²² demonstrates this point. This system uses an ultrathin carbon foil^{1,23} to convert an incident neutral atom into a positive ion. The foil has a thickness of the order of ~100 atoms, allowing transmission down to energies of several hundreds of eV. However, the probability of producing a charged particle decreases dramatically for neutral atom energies below 1 keV. Furthermore, the probability of transmission through the foil also decreases dramatically below about 0.5 keV [e.g., ~25% at 0.2 keV (Ref. 20)], resulting in very low overall efficiencies at energies of a few hundreds of eV.

Graphene, being the strongest material known on Earth, ²⁴ can be made much thinner (as low as 1 atomic layer) than the regular amorphous carbon (~100 atoms thick). Therefore, to first-order, particles passing through few layer graphene (FLG) foils should 1. lose less energy, 2. scatter less, and 3. be transmitted at lower energies than through amorphous carbon foils. The motivation for using graphene foils is obvious: these foils could potentially reduce angular scattering and energy straggling—two effects that can degrade the performance of an instrument.

For ENA imaging, a critical parameter is the charge conversion efficiency of a neutral atom into an ion. Is the charge conversion efficiency of graphene comparable to that of amorphous carbon? The intrinsic differences between the graphene and amorphous carbon should yield different exit charge state distributions. However, the adsorbates—associated with $\sim \! 10^{-7}$ Torr vacuum—are likely to be similar

and drive the exit charge state distributions toward similar values.

In this study, we present measurements of exit charge state distributions of ~ 1 to 50 keV ions passing through graphene foils for the first time, and compare them with measurements using state-of-the-art amorphous carbon foils ($\sim 0.5~\mu g/cm^2$).

A parallel study by Ebert et al.²⁵ presents measurements of ion scattering by the same graphene foils and a comparison with state-of-the-art carbon foils ($\sim 0.5 \ \mu g/cm^2$). In a future study, we plan to characterize the energy straggling and the transmission as a function of energy of ions passing through amorphous carbon and graphene foils.

2 Setup

Our setup design is similar to that of previous works. ^{9,15} The principle is illustrated in Fig. 1(a). A collimated beam of positive ions passes through the foils and scatters. Further downstream, a slit collimates scattered ions in a fan-like shape. Then, deflection plates parallel to the slit separate ions according to their charge. The neutrals go straight through, and positive and negative ions are deflected away from the neutrals in opposite directions. A position sensitive detector measures the particles. Since neutrals and ions of different charges follow different trajectories, they land on the detector at different locations. An image of the impact location shows different bands corresponding to the respective charge states [cf. Fig. 1(b)]. Charge state distributions are readily inferred from the number of counts in each band.

Our imaging detector is a Quantar detector with $\sim 8~\rm cm \times 10~\rm cm$ MCPs in chevron configuration. The active area of the resistive anode is about 4 cm \times 5 cm. We bias the front MCPs at $-100~\rm V$ (to avoid contamination with secondary electrons from the foil) and the resistive anode is at $+2200~\rm V$. All measurements are taken at a pressure of $\sim 2 \times 10^{-7}$ to 8×10^{-7} Torr.

It is possible that the foils have pinholes²⁶ that would let ions pass through without interaction. These ions will show up in the band of positive ions. The pinholes are relatively easy to identify with slow heavy ions²⁷ because the ions that interact with the foil strongly scatter whereas those that pass through the pinholes do not. The ions passing through pinholes in the foil land in a small area on the detector. If pinholes are identified, then we can disregard this small

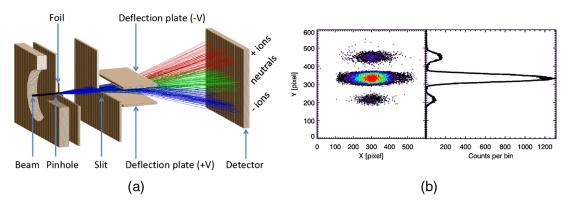


Fig. 1 (a) Exit charge state distribution setup cross section showing the different paths of charged and neutral particles after the foil. (b) Example of an image obtained on the position sensitive detector and its projection onto the vertical axis.

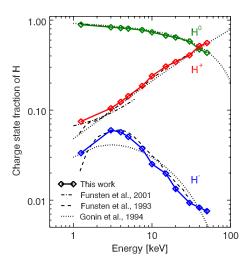


Fig. 2 Charge state fractions of hydrogen after passing an amorphous carbon foil of nominal thickness of $0.5 \mu g/cm^2$. Our results are consistent with the previous work.

area and use counts from neighboring pixels to interpolate the counts of scattered ions in this area.

To mitigate potential systematic differences between the measurements of the positive and negative charge state fractions, we swap polarities of the deflection plates [cf. Fig. 1(a)] and take a second measurement. We calculate the average charge state fraction from both measurements. Note that the difference between the two measurements is usually comparable to the statistical uncertainty of the measurements.

Figure 2 shows the charge state fraction of hydrogen after passing through a regular amorphous carbon foil of nominal thickness $0.5 \mu g/cm^2$. This amorphous carbon foil was purchased from Arizona Carbon Foil Co., Inc, Tucson, Arizona. We compare our results to three previous studies. The curves from Funsten et al. 15,20 are fits through their measurements taken with similar carbon foils. The curves from Gonin et al.²⁶ are from a semiempirical model. The main difference is that the latter study gives charge states as a function of residual energy—the energy of the ions leaving the exit surface after suffering energy loss inside the foil. Although the relative difference between incident and residual energy is small at the higher end of this energy range, it is more significant at the lower end. We do not measure the energy loss with our setup and, thus, we cannot convert the horizontal scale to residual energy at this point. Nevertheless, our results agree very well with the previous published work and that validates our setup and procedure.

3 Charge States at Exit of Graphene Foils

We affixed an FLG (\sim 3 to 7) foils on nickel grids (333 and 1000 lpi). The thickness is determined by Raman spectroscopy using the shape of the two-dimensional peak and by optical transmission. The thickness of the foil ranges from 0.345 nm \times 3 layers to 0.345 nm \times 7 layers. The few layers of graphene are created on copper sheets that are subsequently etched in an acid bath. Once the copper is completely etched away, the bath is neutralized and the graphene foil floating process can take place. The graphene foil is picked up by the grids and removed from the bath. After it is dry, the graphene foil is ready for testing. We also floated two FLG foils (hereafter, abbreviated as

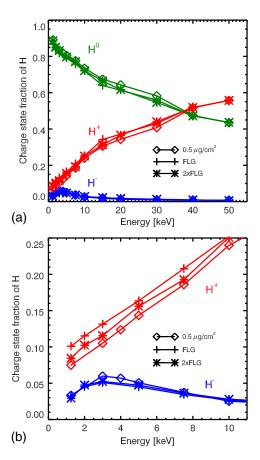
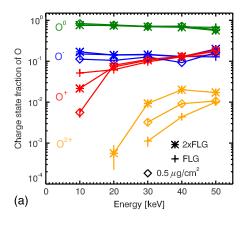


Fig. 3 (a) Charge state fractions of hydrogen after passing through an ultrathin (nominal thickness of 0.5 μ g/cm²) amorphous carbon foil, a single (FLG) graphene foil, and a double (2xFLG) graphene foil. (b) Same data zoomed in. The uncertainties (smaller than the symbols) are derived from propagation of the Poisson errors for the individual measurements.

2xFLG) onto the same substrates in order to improve the quality of the foils and reduce pinholes. The second foil was floated on the top of the first one after it dried. Since (a) the foils are easy to see with the naked eye and (b) we saw that the first foil remained attached during the second floating, we assume that the two foils were successfully stacked together. These foils are still much thinner (\sim 10 times) than a nominal 0.5 μ g/cm² amorphous carbon foil.

Figure 3(a) shows the fraction of neutral, positive, and negative hydrogen exiting graphene and amorphous carbon foils as a function of incident energy. Figure 3(b) shows the same data on expanded scales. The charge state distribution after passing through graphene is roughly similar to that of amorphous carbon. There are absolute differences of up to $\sim 3.7\%$ (at 15 and 30 keV) for positive ions and neutral atoms between the graphene and amorphous carbon foils. The maximum absolute difference for negative ions is <1%. Graphene (FLG and 2xFLG) appears to produce more positive ions than the amorphous carbon at low energies [see Fig. 3(b)].

Figure 4(a) shows the fraction of the different charge states of oxygen after the graphene and amorphous carbon foils. Here again, there are only very small differences between these foils. The largest differences are at the lowest energies or for the minor charge states such as O^+ and O^{2+} .



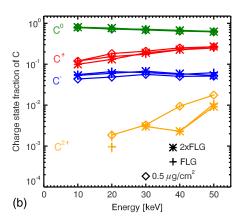


Fig. 4 Same as Fig. 3 for (a) oxygen and (b) carbon.

Graphene produces more negative ions than the amorphous carbon up to $\sim 30 \text{ keV}$.

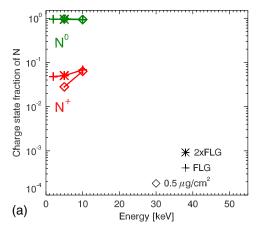
Figure 4(b) shows the fraction of the different charge states of carbon after passing through the same foils. Similarly to hydrogen and oxygen, there is not much difference between the graphene and amorphous carbon foils.

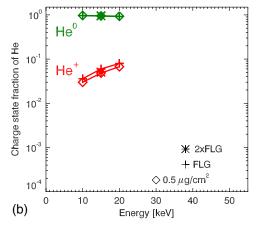
Figure 5 shows the fraction of the different charge states of nitrogen, helium, and argon after passing through graphene and amorphous carbon. There are only very small differences between the two foil types. For argon, there is no systematic difference.

4 Discussion

We measured the charge state fractions of ions after passing through graphene and amorphous carbon foils. The measurements cover an energy range of ~ 1 to 50 keV. We find that:

- The exit charge state distributions of incident ions used in this study transiting ultrathin foils of graphene and amorphous carbon are similar. Some small differences are listed below.
- Graphene produces slightly more positive ions after the foil than the amorphous carbon for hydrogen.
- Graphene produces slightly more negative ions than the amorphous carbon for oxygen and carbon.
- There are no systematic differences of the charge state fractions after graphene or amorphous carbon for argon.





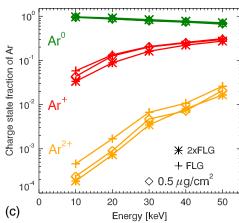


Fig. 5 Same as Fig. 3 for (a) nitrogen, (b) helium, and (c) argon.

The similarity of the exit charge distribution may indicate that the adsorbate layer at the exit surface that is assumed to be similar for both types of foils may play an important role in defining the exit charge state distribution.

The charge state of a particle exiting a foil in our energy range is determined mainly by the three factors: 1. the particle atomic number, 2. the foil surface composition, and 3. the particle velocity when it leaves the surface. In our experiment, it is reasonable to assume that the particle velocity is roughly equal for graphene and amorphous carbon even though graphene is supposed to be much thinner. This

argument is supported by our data in the sense that there is no systematic difference that applies to all species. Even if this assumption was not true, the difference would be small because the total energy loss is a small fraction (a few to $\sim 15\%$) of the total energy.

The fact that both graphene and amorphous carbon foils yield similar exit charge state distributions suggests that the common adsorbates material is perhaps the governing driver for the exit charge state distribution. It is well known that amorphous carbon foils have layers of impurities^{28,29} on their surface that can account for a significant fraction of the total foil thickness. ^{23,30} Thus, it is likely that the graphene foils have similar impurities sticking to their surface. The chemistry might be different, but graphene, similarly to graphite, adsorbs, and desorbs various molecules.

Measurements of angular scattering distributions with graphene foils by Ebert et al.²⁵ support the hypothesis of the impurities on the surface of the foil. We mentioned in Sec. 1 a future study to measure the energy straggling and the transmission as a function of energy of ions passing through amorphous carbon and graphene foils. The measurements presented here, the energy straggling, and the angular scattering measurements can provide important clues to determine whether graphene foils also have layers of impurities.

This study shows that graphene performs similarly to amorphous carbon when it comes to converting ENA into ions. Thus, it is conceivable that graphene foils could replace amorphous carbon foils in ENA detectors in the future, provided that they present an advantage for other properties such as angular scattering and energy loss.

Acknowledgments

The authors would like to acknowledge the contributions from Jessica Armstrong, Dave Cronk, Guy Grubbs, Dave McComas, Greg Miller, Ed Patrick, Steve Petrinec, Amanda Richter, Ben Rodriguez, Ken Smith, and Phil Valek. This research was funded by an internal research and development grant at Southwest Research Institute and internal research funding at Lockheed Martin Space Systems Company, Advanced Technology Center.

References

- D. J. McComas et al., "Ultra-thin (~10 nm) carbon foils in space instrumentation," Rev. Sci. Instrum. 75(11), 4863–4870 (2004).
- 2. S. M. Ritzau and R. A. Baragiola, "Electron emission from carbon foils induced by keV ions," Phys. Rev. B 58(5), 2529-2538 (1998).
- F. Allegrini et al., "Determination of low-energy ion-induced electron yields from thin carbon foils," Nucl. Instrum. Methods B 211(4), 487-494 (2003).
- 4. E. A. Kurz, "Channel electron multipliers," in *American Laboratory*, International Scientific Communications, Frederick I. Scott, Ed., Fairfield, Connecticut (1979)
- J. L. Wiza, "Microchannel plate detectors," *Nucl. Instrum. Methods A* **162**(1–3), 587–601 (1979).
- D. Hovestadt et al., "CELIAS-charge, element and isotope analysis system for SOHO," *Solar Phys.* 162(1-2), 441–481 (1995).
 D. T. Young et al., "Cassini plasma spectrometer investigation," *Space Sci. Rev.* 114(1-4), 1-112 (2004).
- A. Bürgi et al., "Charge exchange of low-energy ions in thin carbon foils," *J. Appl. Phys.* **68**(6), 2547–2554 (1990).

 A. Bürgi et al., "Charge exchange of low energy ions in thin carbon
- foils. II. Results for ions of B, C, F, Ne, Na, Si, S, Cl, Ar, K, and Fe," J. ol. Phys. 73(9), 4130-4139 (1993).
- 10. M. Gonin, R. Kallenbach, and P. Bochsler, "Charge exchange of hydrogen atoms in carbon foils at 0.4-120 keV," Rev. Sci. Instrum. **65**(3), 648–652 (1994).

- 11. M. Gonin, R. Kallenbach, and P. Bochsler, "Charge exchange of atoms with high first ionization potentials in carbon foils at $E \le 35 \text{ keV/u}$," *Nucl. Instrum. Methods B* **94**(1–2), 15–21 (1994).
- 12. M. Gonin et al., "Charge exchange of low energy particles passing through thin carbon foils: dependence on foil thickness and charge state yields of Mg, Ca, Ti, Cr and Ni," Nucl. Instrum. Methods B **101**(4), 313–320 (1995)
- 13. R. Kallenbach et al., "Charge exchange of hydrogen ions in carbon foils," *Nucl. Instrum. Methods B* **83**(1–2), 68–72 (1993). 14. R. Kallenbach et al., "Charge exchange of B, C, O, Al, Si, S, F and Cl
- passing through thin carbon foils at low energies: formation of negative ions," *Nucl. Instrum. Methods B* **103**(2), 111–116 (1995).
- 15. H. O. Funsten, D. J. McComas, and B. L. Barraclough, "Ultrathin foils used for low-energy neutral atom imaging of the terrestrial magneto-sphere," *Opt. Eng.* **32**(12), 3090–3095 (1993).

 16. H. O. Funsten, B. L. Barraclough, and D. J. McComas, "Shell effects observed in exit charge state distribution of 1–30 keV atomic projec-
- tiles transiting ultrathin carbon foils," Nucl. Instrum. Methods B 80,
- 17. H. O. Funsten, B. L. Barraclough, and D. J. McComas, "Interactions of slow H, H2, and H3 with thin carbon foils," Nucl. Instrum. Methods B **90**(1–4), 24–28 (1994)
- 18. H. O. Funsten, "Formation and survival of H- and C- ions transiting ultrathin carbon foils at keV energies," Phys. Rev. B 52(12), 8703-
- 19. H. O. Funsten, S. M. Ritzau, and R. W. Harper, "Negative helium ions exiting a carbon foil at keV energies," Phys. Rev. B 63(15), 155416
- 20. H. O. Funsten, D. J. McComas, and M. Gruntman, "Energetic neutral atom imaging of the outer heliosphere-LISM interaction region," in The Outer Heliosphere: The Next Frontiers, K. Scherer et al., Eds., COSPAR Colloquia Series, Vol. 11, p. 237, Pergamon, Amsterdam
- 21. D. J. McComas et al., "IBEX—interstellar boundary explorer," Space
- Sci. Rev. 146(1-4), 11-33 (2009).

 22. H. O. Funsten et al., "The interstellar boundary explorer high energy (IBEX-Hi) neutral atom imager," Space Sci. Rev. 146(1-4), 75-103
- 23. F. Allegrini et al., "Energy loss of 1-50 keV H, He, C, N, O, Ne, Ar ions transmitted through thin carbon foils," *Rev. Sci. Instrum.* 77(4),
- 044501 (2006).

 24. C. Lee et al., "Measurement of the elastic properties and intrinsic strength of monolayer graphene," *Science* 321(5887), 385–388 (2008).
- 25. R. W. Ebert et al., "Angular scattering of 1-50 keV ions through graphene and thin carbon foils: potential applications for time-of-flight ion mass spectrometers and energetic neutral atom sensors in space," in press (2014).
- 26. H. O. Funsten, B. L. Barraclough, and D. J. McComas, "Pinhole detection in thin foils used in space plasma diagnostic instrumentation, *Rev. Sci. Instrum.* **63**(10), 4741–4743 (1992).
- 27. H. O. Funsten, D. J. McComas, and B. L. Barraclough, "Thickness
- flight technique. A new approach to secondary ion mass spectrometry," *Rev. Sci. Instrum.* **60**(10), 3188–3194 (1989).

 30. G. Both et al., "Ultrathin foils for Coulomb explosion experiments,"
- Rev. Sci. Instrum. 58(3), 424–427 (1987).

Frédéric Allegrini is a principal scientist at the Southwest Research Institute. He received a MS in physics from the University of Lausanne and a PhD in space physics from the University of Berne. His primary interests are space plasma instrumentation, carbon foils for space applications, and data analysis in the area of heliospheric physics. He joined the Department of Physics and Astronomy at UTSA as an adjoint assistant professor.

Robert W. Ebert is a research scientist at the Southwest Research Institute. He has earned degrees from the University of Calgary (BS in astrophysics), the University of Toronto (MS in engineering physics), and the University of Texas at San Antonio (PhD in space physics). His research interests include the solar wind and its interaction with Jupiter's magnetosphere, the origin, acceleration, and transport of energetic particles in interplanetary space, and developing new technologies for space science instrumentation.

Stephen A. Fuselier is the director for Space Science, Space Science and Engineering Division, Southwest Research Institute. He has more than 25 years of experience as a scientist, project manager, and senior manager for scientists, engineers, and technicians. He is a co-investigator and sensor lead on the Interstellar Boundary

Explorer (IBEX) mission, co-investigator and lead for the Hot Plasma Composition Experiment (HPCA) on the Magnetospheric Multiscale, and co-investigator on IMAGE, and ROSINA.

Georgios Nicolaou is a PhD candidate in University of Texas at San Antonio (UTSA) and he is working as research assistant at Southwest Research Institute. He received MS in physics from the National and Kapodistrian University of Athens. His main interests are data analysis in the area of planetary and heliospheric physics, thermodynamics of space plasma and carbon foils for space instrumentation applications.

Peter Bedworth is a senior staff chemist at Lockheed Martin Space Systems Advanced Technology Center. He received his BS (Chemistry) from Augusta College, his PhD (Chemistry) from University of South Carolina under James Tour and was a research fellow at Caltech with Seth Marder. He has co-authored numerous

patents and publications. His primary research thrust is the synthesis and characterization of 2D and structured materials for optical, electronic, magnetic and membrane devices.

Steve Sinton is a Lockheed Martin Fellow at the Advanced Technology Center of LM Space Systems Company. He received a PhD in physical Chemistry from the University of California, Berkeley. His primary interests are materials science and nanotechnology in aerospace and commercial applications.

Karlheinz J. Trattner received a BSc in physics from the University of Graz, Austria, in 1983; a MSc in geophysics from the University of Graz, Austria, in 1989; and a PhD in space science from the University of Graz, Austria in 1992. He has held a post at Lockheed Martin ATC in Palo Alto California for 16 year and worked as a Science Lead on several investigations about solar-terrestrial interactions and magnetic reconnection.